A VO service for the European VLBI Network

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ESCAPE WP4 Technology Forum 2

JIVE

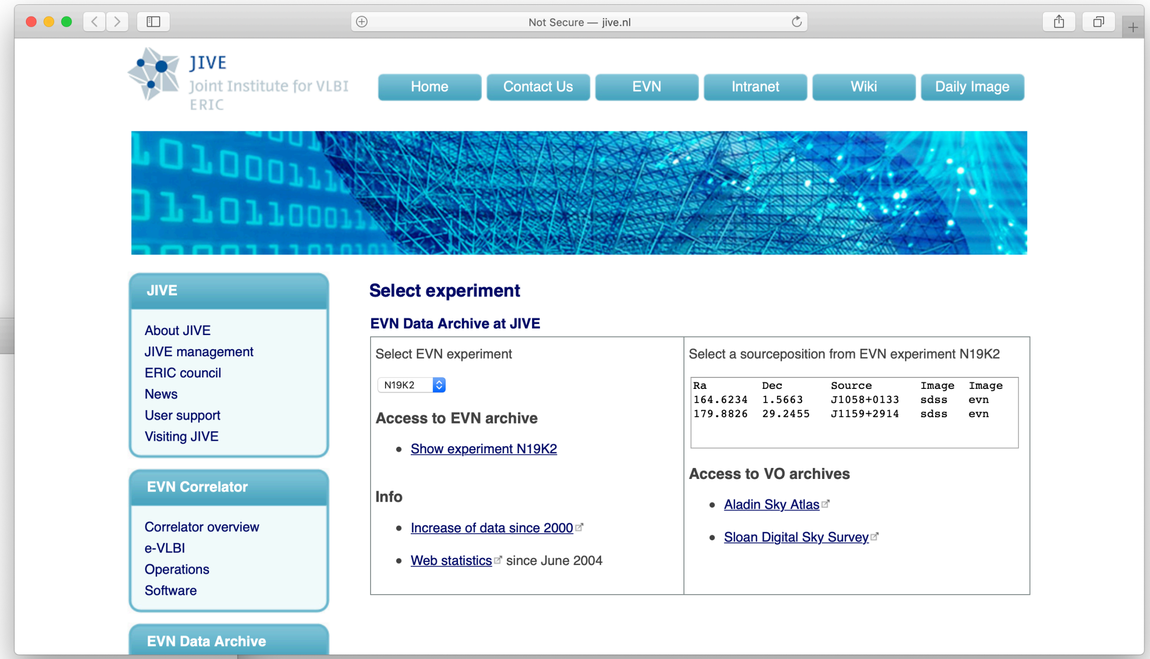
Joint Institute for VLBI ERIC

ESCAPE - The European Science Cluster of Astronomy & Particle Physics ESFRI Research Infrastructures has received funding from the European Union's Horizon 2020 research and innovation programme under the Grant Agreement n° 824064.

EVN & JIVE

- EVN: European VLBI Network
  - Collaboration between radio observatories in Europe and beyond (South-Africa, Puerto-Rico, China, Korea)
  - Heterogeneous array
  - PI driven
- JIVE: Joint Instutute for VLBI ERIC
  - Support institute for the EVN
  - Operates the EVN correlator and hosts the EVN data archive

EVN Archive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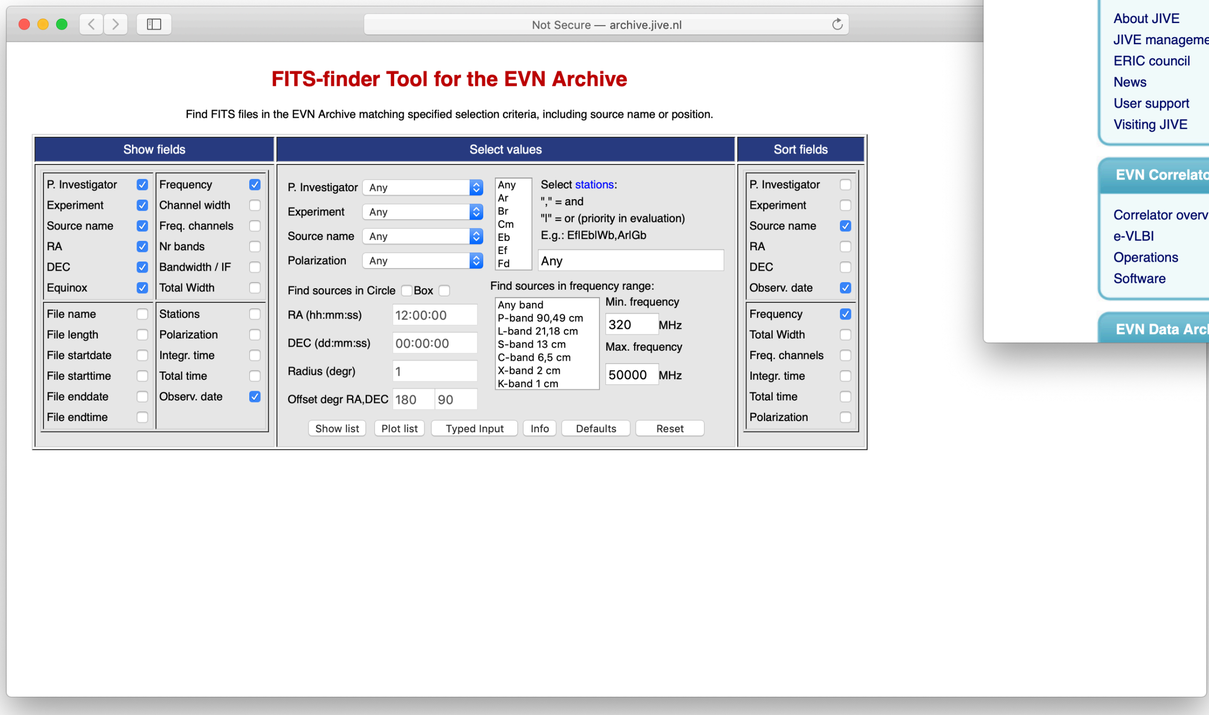
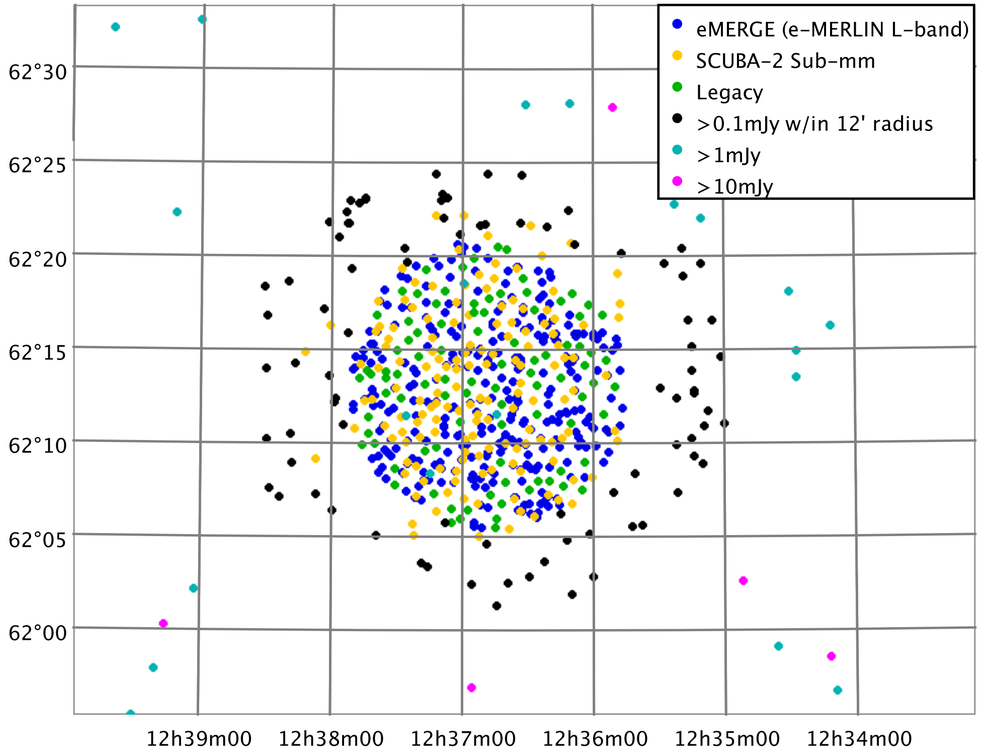
N19K2
Access to EVN archive
Show experiment N19K2
Info
Increase of data since 2000
Web statistics since June 2004
Select a sourceposition from EVN experiment N19K2
Ra Dec Source Image Image
164.6234 1.5663 J1058+0133 sdss evn
179.8826 29.2455 J1159+2914 sdss evn
Access to VO archives
Aladin Sky Atlas
Sloan Digital Sky Survey


Not Secure — archive.jive.nl
FITS-finder Tool for the EVN Archive
Find FITS files in the EVN Archive matching specified selection criteria, including source name or position.
Show fields
Select values
Sort fields
P. Investigator
Experiment
Source name
RA
DEC
Equinox
Frequency
Channel width
Freq. channels
Nr bands
Bandwidth / IF
Total Width
File name
File length
File startdate
File starttime
File enddate
File endtime
Stations
Polarization
Integr. time
Total time
Observ. date
Any
Select stations:
"," = and
"|" = or (priority in evaluation)
E.g.: Ef|Eb|Wb,Ar|Gb
Find sources in Circle
Box
RA (hh:mm:ss)
12:00:00
DEC (dd:mm:ss)
00:00:00
Radius (degr)
1
Offset degr RA,DEC
180
90
Find sources in frequency range:
Any band
P-band 90,49 cm
L-band 21,18 cm
S-band 13 cm
C-band 6,5 cm
X-band 2 cm
K-band 1 cm
Min. frequency
320
MHz
Max. frequency
50000
MHz
Show list
Plot list
Typed Input
Info
Defaults
Reset

Data products

- Visibility data (“UV data”; FITS-IDI)
  - No in-beam calibrators -> Multiple sources per observation
  - Continuum and spectral line data
  - Pulsar observations: multiple bins
  - MPC observations: multiple field centers
- Filterbank data (Pulsars, FRBs; PSR-FITS)
  - Time-series
- Calibration data
  - Flagging, amplitude calibration, observation schedule, observation logs
- Diagnostic plots (from pipeline)

Multiple Phase Centers



Radcliffe et. al.

699 sources in GOODS-N

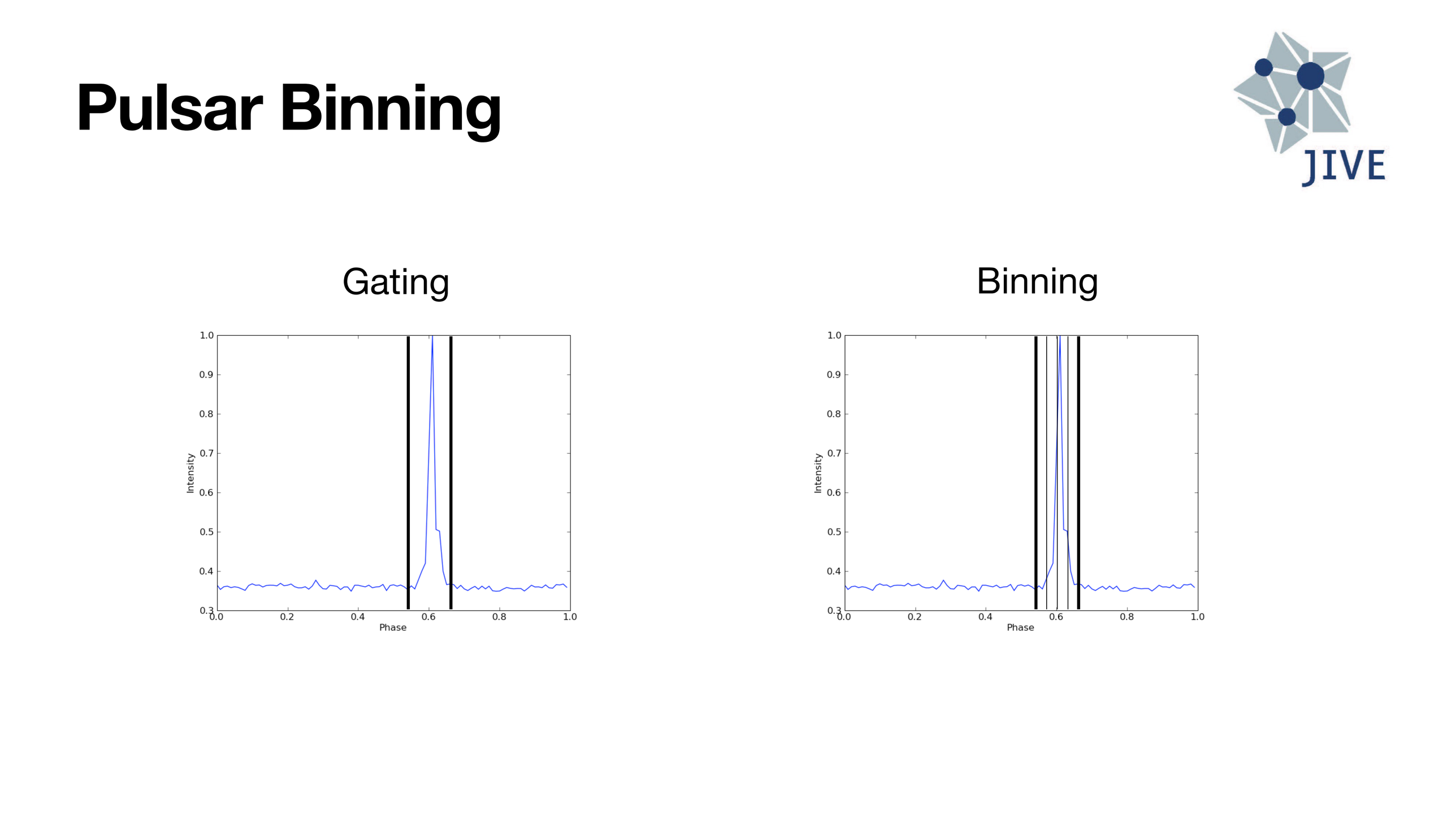
Two areas:

- 15’ central area
- 20’ outer annulus

Multi-source Self Calibration

arXiv:1601.04452

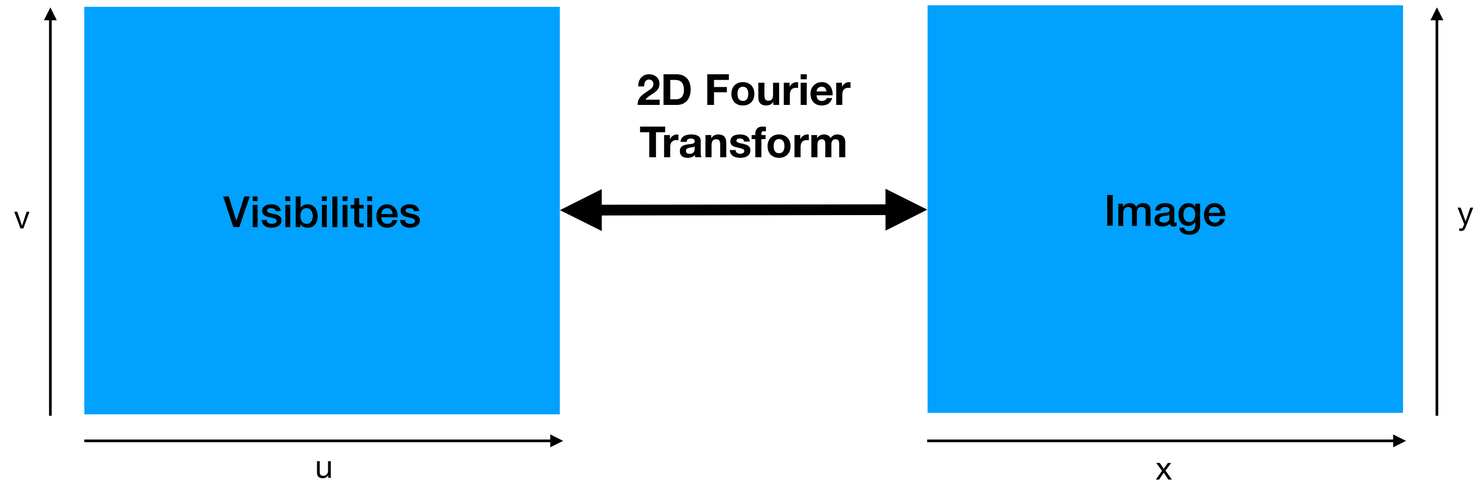
Pulsar Binning



Gating

Binning

Visibilities & Images



- UV plane is not completely filled
- Visibilities are (partly) self-calibrated
- Imaging algorithm choices depend on scientific goal

VO use case

- Access historic data for high-resolution follow-up of:
  - Gravitational Wave events
  - Gamma Ray Bursts
  - Fast Radio Bursts

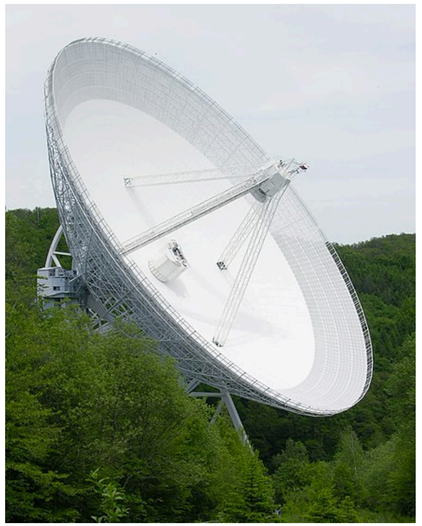
- Standardized accès to archive data for science platform
  - JupiterLan environment; see presentation by Aard Keimpema tomorrow

VO protocols that match: ObsTAP and Datalink

Implementation using DaCHS

- Why DaCHS?
  - Some in-house knowledge (at ASTRON)
  - Python
  - Visit by Marcus (to ASTRON)
  - Implements TAP and Datalink service
- DaCHS runs alongside existing Archive interface
  - Linking to data products in Existing archive
- DaCHS ingests CSV data generated by separate Python “fitscrawler” Tool
  - FITS-IDI stores Important metadata in (large) binary tables

ObsCore representation of visibility data



- Spatial extent determined by several factors:
  - FoV of individual telescopes (in particular the largest telescope)
  - Time and frequency smearing
  - Projected longest baseline (distance between telescopes)
- Approximated assuming maximal amplitude loss of 50% and ignoring projection effects.
- Problem: Antenna diameters are not stored in FITS-IDI files
- Problem: Antenna names are not always unique

ObsCore representation of visibility data

- s_resolution based on longest baseline
  - Approximation; should be based on synthesised beam (from UV coverage)
- t_resolution is the correlator integration time
- t_exptime is calculated by summing integration time
  - Each source in the observation becomes separate ObsCore dataset
- em_min and em_max calculated based on minimum and maximum observed frequency
  - Dual S/X band observations should probably be split into separate ObsCore datasets

ObsCore representation of visibility data

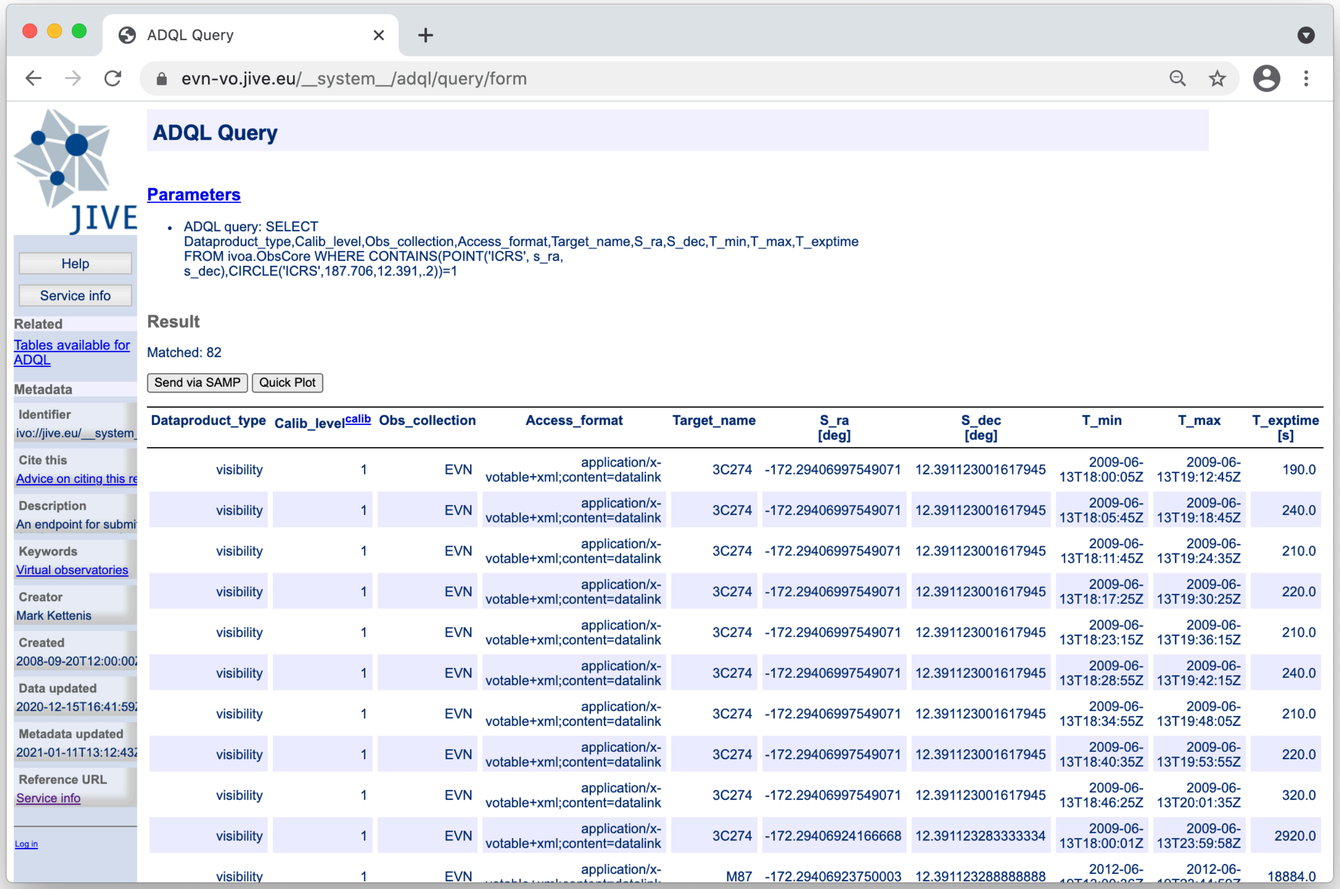
- Some observations are correlated multiple times with different parameters
  - “continuum” and “spectral line” get its own ObsCore dataset
- MPC correlations result in multiple sets of output file
  - Each phase centre gets its own ObsCore dataset
- Pulsar binning/gating
  - Each bin gets its own ObsCore dataset (including “off-pulse” bin)
- access_url is a Datalink

Datalink for FITS-IDI

- Single observation split into several FITS-IDI files of ~2GB
  - Return Datalink table with rows for each chunk
- Calibration data will be added in the future
- Considering adding pipeline images as previews
  - These are often very rough!
- Considering adding diagnostic plots as secondary datalink

Preliminary service

- TAP interface URL: https://evn-vo.jive.eu/tap
- Covers almost all public data in EVN Archive
- Almost 23K ObsCore datasets (for 150 TB of data)
  - From ~2000 observations
- Global metadata for the service still needs to be validated
- Not registered yet



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visibility	1	EVN	application/x-votable+xml;content=datalink	3C274	-172.29406997549071	12.391123001617945	2009-06-13T18:05:45Z	2009-06-13T19:18:45Z	240.0
visibility	1	EVN	application/x-votable+xml;content=datalink	3C274	-172.29406997549071	12.391123001617945	2009-06-13T18:11:45Z	2009-06-13T19:24:35Z	210.0
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visibility	1	EVN	application/x-	M87	-172.29406923750003	12.391123288888888	2012-06-	2012-06-	18884.0

Some initial feedback

- em_min, em_max and em_res_power given in wavelengths
  - This is unatural for radio astronomy
  - Posible solution: add f_min, f_max and f_resolution?

- Datalink acess_url initially surprises users

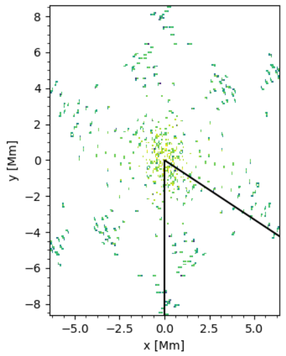
- Some records contain metadata that is wrong
  - Early FITS-IDI files are sometimes incorrect
  - Archive contains FITS-IDI files with WSRT local interferometry data that probably should be removed

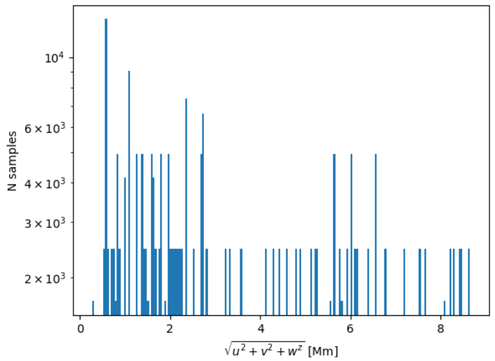
Extending ObsCore

- Ongoing discussion within the IVOA radio IG of common extensions:
  - Characterization of UV coverage (eccentricity, filling factor)
  - Largest spatial scale
  - Frequency-based characterisation of observed spectrum

- Plan to add DataCite DOIs for each observation in the EVN archive
  - Which ObsCore field should be used?

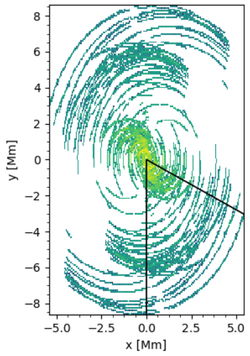
UV space characterization

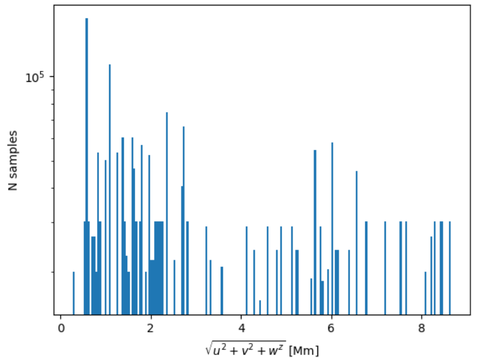




$e = 0.74$
$f = 0.03$

calibrator source





$e = 0.63$
$f = 0.25$

target source

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